

1 **A model study of tropical gravity wave changes with El Niño and La Niña**
2 **conditions and wave forcing of the quasibiennial oscillation**

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ABSTRACT

14 Using an idealized model framework with high-frequency tropical latent
15 heating variability derived from global satellite observations of precipitation
16 and clouds, we examine the properties and effects of gravity waves in the
17 lower stratosphere, contrasting conditions in an El Niño and La Niña year.
18 The model generates a broad spectrum of tropical waves including planetary-
19 scale waves through mesoscale gravity waves. We compare modeled monthly-
20 mean regional variations in wind and temperature with reanalyses, and we val-
21 idate the modeled gravity waves using satellite- and balloon-based estimates
22 of gravity wave momentum flux. Some interesting El Niño Southern Oscilla-
23 tion (ENSO) changes in the gravity spectrum of momentum flux are found in
24 the model which are discussed in terms of the ENSO variations in clouds, pre-
25 cipitation, and large-scale winds. The modeled intermittency in gravity wave
26 momentum flux is shown to be very realistic compared to observations, and
27 the largest amplitude waves are related to significant gravity wave drag forces
28 in the lowermost stratosphere. This strong intermittency is generally absent
29 or weak in climate models due to deficiencies in parameterizations of gravity
30 wave intermittency. Our results suggest a way forward to improve model rep-
31 resentations of lowermost stratospheric quasi-biennial oscillation winds and
32 teleconnections.

33 1. Introduction

34 Long-range weather forecast skill has demonstrated links to the tropical stratosphere. The quasi-
35 biennial oscillation (QBO), a reversal of the tropical lower stratosphere zonal mean winds roughly
36 every other year, is an important source of predictability for seasonal forecasts of the North At-
37 lantic Oscillation (NAO) [Scaife et al. 2014a]. Although confined to tropical latitudes, the QBO
38 has clear connections to polar stratospheric extreme vortex events [Holton and Tan 1980] and ex-
39 tratropical surface weather conditions [Thompson et al. 2002; Garfinkel et al. 2012]. The QBO
40 also modulates the chemical composition of the stratosphere, for example dominating interannual
41 variability in tropical stratospheric water vapor with associated effects on chemistry and tempera-
42 ture [Mote et al. 1996; Randel et al. 2004].

43 The key characteristics of the QBO are zonal mean winds that oscillate from easterly to westerly
44 with an average period of 28 months and ranging from ~ 22 to 34 months. The circulation is not
45 locked to the annual cycle, rather the period is inversely related to atmospheric wave momentum
46 transport, or more specifically to the divergence of Eliassen-Palm flux [Dunkerton 1997], which is
47 often called wave drag or wave forcing. A range of studies have identified tropical gravity waves
48 as a crucial component of the momentum transport necessary to drive the QBO, with current
49 estimates suggesting more than half of the flux is carried by gravity waves that are generated
50 by tropical convection (e.g. Kawatani et al. [2010a]). Global characterizations of gravity waves
51 and their momentum transport remain an observational challenge due to their small scales, high
52 frequencies, and intermittent occurrences [Geller et al. 2013; Alexander 2015]. Models therefore
53 play a major role in our understanding of tropical gravity waves and their effects on circulation.

54 Despite the fact that the basic wave-mean flow interaction mechanism that forces the QBO has
55 been understood for decades, most climate models still do not have a QBO. The global climate

56 and weather forecasting models that do include a QBO rely on parameterization of sub-grid-scale
57 gravity wave forcing. Such parameterizations often assume an average set of wave properties,
58 continually forced at all times and all longitudes. A few climate models include varying gravity
59 wave properties that are tied to the model’s parameterized convection (e.g. Richter et al. [2014],
60 Bushell et al. [2015]), but large uncertainties in specifying the properties of the sub-gridscale
61 gravity waves remain [Schirber et al. 2015].

62 In this study, we employ a global model that is uniquely constrained by observations in multiple
63 ways in order to examine interannual, subseasonal, and geographical variability in tropical gravity
64 waves and their effects on the stratospheric circulation. The model is designed to represent the
65 scales of gravity waves that are observed by limb-sounding satellite measurements, which provide
66 constraints for the modeled gravity wave momentum transport. Global precipitation and cloud
67 observations provide constraints on variability in latent heating that force the waves in the model.
68 Finally, zonal mean winds in the model are relaxed to those in reanalysis, and the model is initial-
69 ized with reanalyzed zonal mean temperature in order to realistically constrain wave propagation
70 and interactions with the circulation.

71 In section 2, we describe the precipitation and cloud data that will be used to estimate latent
72 heating that forces tropical waves in the model. Section 3 describes the model experiment design,
73 the latent heating algorithm, and the properties of tropical clouds, precipitation, and latent heating
74 for the two experiments in El Niño and La Niña conditions. Section 4 describes the model results
75 and compares them to observations, including tropical tropopause wind and temperature varia-
76 tions, the properties of the gravity waves, and wave driving of the QBO. Section 5 is a discussion
77 of the results and their implications for global climate and weather modeling, and section 6 is a
78 summary with conclusions.

79 2. Data Description

80 *a. Background*

81 The launch of the Tropical Rainfall Measuring Mission (TRMM) satellite in 1998 began an
82 era of high-resolution global precipitation measurement. TRMM products include the TMPA
83 (3B42) gridded rain rates at $0.25^\circ \times 0.25^\circ$ spatial resolution and 3-hourly time intervals [Huffman
84 et al. 2007]. Ryu et al. [2011] described an algorithm for computing the 3-dimensional time-
85 dependent latent heating field using TRMM 3B42 rain rates, and ancillary cloud-top height derived
86 from global merged geostationary satellite observations of infrared brightness temperatures. The
87 resulting three-dimensional time-dependent latent heating field defined sources for waves with
88 periods longer than 6 hours in a dry dynamical primitive equation model. Ortland et al. [2011]
89 used this method to study the sensitivity of the wave Eliassen-Palm flux (EP-flux) spectrum to
90 spatial resolution in the model. They showed that at the 3-hourly temporal resolution of the TRMM
91 3B42 data, spatial resolutions resolving wavenumbers higher than 60 gave only minor increases in
92 the EP-flux. Thus for modeling gravity waves, the 3-hourly time resolution was clearly a limiting
93 factor. Higher-frequency gridded precipitation products available include CMORPH [Joyce et
94 al. 2004] with 8-km spatial and 30-min temporal resolution, GSMAP with 0.1° spatial and 1-
95 hourly temporal resolution [Ushio et al. 2009], and a recent product called IMERG [Huffman et
96 al. 2015, ATBD] with 0.1° spatial and 30-min temporal resolution based on new measurements
97 from the Global Precipitation Missions (GPM). IMERG is currently only available beginning in
98 2014, although there are plans to extend the IMERG record back in time. For the present study, we
99 seek to model the historical period 2006-2007, a time when we have high-resolution satellite limb-
100 sounding observations that we can use for validation of gravity wave momentum fluxes [Alexander
101 2015]. We therefore utilize CMORPH rain rates degraded to $0.25^\circ \times 0.25^\circ$ spatial resolution but

utilizing the full 30-min temporal resolution. The focus in this work will be on tropical convective wave sources within $\pm 30^\circ$ of the equator.

b. CMORPH Precipitation

Like other high-resolution precipitation products, CMORPH takes advantage of the frequent sampling of geostationary infrared measurements and combines these with higher-quality microwave precipitation measurements to create a gridded spatio-temporally varying precipitation rate product. Sapiano and Arkin [2009] found 3-hourly CMORPH $0.25^\circ \times 0.25^\circ$ data showed the highest correlations against gauge data among several comparable data sets. However, all showed a tendency to underestimate rainfall over the tropical Pacific Ocean. Habib et al. [2012] evaluated CMORPH at $8 \text{ km} \times 8 \text{ km}$, 30-min resolution against dense rain gauge observations and radar-based estimates in Louisiana. CMORPH was found to have negligible bias over the 28-day study period, high detection skills and rainfall occurrence distributions that compare very well to the radar. Significant biases occurred on event scales, and missed and false-rain detections were $\sim 20\%$, but these errors are reduced considerably through aggregation. Our degradation to $0.25^\circ \times 0.25^\circ$ resolution partly serves that purpose, and we will further focus on statistics of wave generation with a full month of data.

c. Global-merged Infrared Brightness Temperature and Cloud-top Height

TRMM ancillary data include global merged infrared brightness temperatures utilizing the international set of geostationary measurements. The merged data is available equatorward of 60° at 4-km spatial resolution and 30-min time resolution. As for the rain rate data, we degrade these brightness temperatures to $0.25^\circ \times 0.25^\circ$ spatial resolution but retain the 30-min temporal resolution. Occasional gaps in coverage occur but are filled by linear interpolation in time utilizing

124 measurements in the previous and following 1.5 hours to achieve a continuous coverage. A small
125 but persistent gap in coverage occurs in the tropical Southern Hemisphere eastern Pacific, but
126 since this is a generally dry region, it is unlikely to affect our results. Cloud-top height is es-
127 timated by matching the observed brightness temperature to regional temperature profiles taken
128 from the MERRA reanalysis [Rienecker et al. 2011].

129 Figure 1 shows sample snapshots of the resulting rain rates and cloud-top heights at
130 $0.25^\circ \times 0.25^\circ$, 30-min resolution.

131 **3. Experiment Design**

132 *a. Latent heating*

133 Using the precipitation rates and cloud top heights from section 2, a three-dimensional, time-
134 dependent tropical latent heating field is calculated with the algorithm described in detail in Ryu
135 et al. [2011]. Briefly summarizing the procedure, a combined rain-rate and cloud-height criterion
136 categorizes the heating profile as convective or stratiform type. For the convective-type rain, heat-
137 ing is specified as positive everywhere with a half-sine vertical profile shape, and heating depth
138 and peak heating are functions of rain-rate and cloud top height. For stratiform-type rain, melting
139 level is a function of rain rate, and peak upper tropospheric heating rate and peak lower tropo-
140 spheric cooling rate are functions of rain-rate and cloud height. The algorithm is similar to the
141 TRMM Spectral Latent Heating (SLH) product [Shige et al. 2007]. In particular, it includes a
142 factor that accounts for horizontal transport of precipitation from convective to stratiform regions.

143 This three-dimensional latent heating is computed at the same 30-min resolution as the cloud
144 and precipitation data. The latent heating field can be compared to other existing latent heating

145 products after sufficient averaging. The appendix provides some additional information about the
146 latent heating.

147 *b. Model*

148 A global nonlinear spectral model is used to simulate waves forced by the space-time-varying
149 tropical latent heating. The model was described previously in Ortland et al. [2011]. In the present
150 cases, the model is initialized with November monthly mean zonal winds and temperatures and
151 run for two months, through the end of December. We subsequently analyze the December re-
152 sults only. The model troposphere is forced with the three-dimensional, time-varying latent heat-
153 ing field, which forces a broad spectrum of waves, including global-scale equatorial Rossby and
154 Kelvin waves through mesoscale gravity waves. Newtonian cooling is applied to the perturbation
155 temperature, where a perturbation is defined as the deviation from the daily MERRA zonal mean.
156 The time scale of the Newtonian cooling is 5 days in the stratosphere and 25 days in the middle
157 and upper troposphere [Ryu et al. 2011]. In the lower troposphere below 650 hPa ($\sigma > 0.7$), the
158 Newtonian cooling time scale gradually decreases to 8 days at the lowest model level [Held and
159 Suarez 1994]. The zonal-mean zonal wind is relaxed to a time-varying state defined by the daily
160 zonal-mean MERRA zonal wind at a rate of $.05 \text{ d}^{-1}$ to ensure realistic QBO shear throughout
161 the simulation. The vertical grid consists of 130 sigma levels with a uniform resolution of 500 m
162 between the surface and 65 km. The top ~ 30 km of the model serves as a sponge layer where
163 wind perturbations are strongly damped with Rayleigh friction. The friction starts from zero at 30
164 km and ramps with with a tanh function shape to a maximum of 7 d^{-1} above 45 km. An eighth-
165 order horizontal hyper-diffusion is used to prevent energy from accumulating at the smallest model
166 scales. It is set to damp the smallest scale at a rate of 1 d^{-1} . The horizontal resolution is truncated
167 at T120, resulting in a resolution of ~ 150 km, a scale chosen to permit simulation of gravity waves

resolved in limb-sounding satellite observations [Alexander 2015] that we will use to validate the modeled gravity waves. We force the model with the zonally-symmetric component of the heating removed, and focus on analysis of waves with periods shorter than 30 days.

c. Experiments

We simulate two December periods in 2006 and 2007. These represent, respectively, weak El Niño and strong La Niña events as seen in the time series of the ENSO 3.4 index shown in Figure 2. Following the characteristic pattern for such conditions, precipitation in the December 2006 period was strongest in the central and eastern Pacific, while the precipitation maximum shifted to the Indian Ocean/Maritime Continent region in December 2007. Both Decembers also included significant precipitation variability associated with the Madden-Julian Oscillation (MJO), also shown in Fig. 2. The MJO is only active during the last 13 days of December 2006, while it is active throughout December 2007. For 2006, the MJO activity is focused in the eastern Indian Ocean and Maritime Continent (phases 3-4), whereas in 2007 the signal propagates from Africa to the Maritime Continent (phases 1-4).

The properties of waves forced by convective heating are sensitive to both the strength of the heating and the depth of the heating (e.g. Salby and Garcia [1987]; Bergman and Salby [1994]; Holton et al. [2002]; Beres [2004]). Geller et al. [2016] suggest that changes in these parameters with ENSO may explain previously observed sensitivity of the QBO to ENSO conditions. Thus to help place our wave analysis results presented later in context, we compare here occurrence frequency distributions of rain rate and cloud-top height for our two cases.

Figure 3 compares distributions of tropical rain rates and cloud top heights for December 2006 and December 2007 at latitudes most relevant to the QBO (10°S - 10°N). These distributions represent “convective” pixels only, defined as those with rain rates greater than 1.6 mm hr^{-1} . The two

distributions are very similar, with no significant differences in the mean values. The La Niña year (2007) displays slightly more of the deepest clouds and higher rain rates, however there are fewer convective pixels overall, only 3.5% in December 2007 compared to 4.3% in December 2006. While ENSO dramatically affects the geographical patterns in precipitation, Fig. 3 suggests that globally these statistics of tropical convective rain rate and cloud depth remain relatively constant.

4. Results

a. Monthly-mean Patterns in Latent Heating and Tropopause Wind and Temperature

Figure 4 shows the December mean heating rates at the 400 hPa level at $0.25^\circ \times 0.25^\circ$ resolution. Characteristic patterns associated with El Niño (2006) and La Niña (2007) are clearly visible in these panels. Note that rates in Fig. 4 are shown at finer horizontal resolution than the model. To force the model, 30-min heating rates are spectrally decomposed with spatial spherical harmonics and truncated to T120 ($\sim 1.5^\circ$ resolution), and the 30-min rates are linearly interpolated in time to the model 3-min timestep.

Figure 5 compares December-mean tropopause (~ 17 km) temperature maps for the MERRA reanalysis and the model. The deepest latent heating in the model extends only to the 15-km level, so these maps represent the dynamical responses to latent heating above those levels that are directly forced. Note that the model initial conditions are zonally symmetric, so the longitudinal variations in temperature result solely from wave responses to the tropospheric heating. The top panels compare December 2006, and December 2007 is compared in the bottom panels. The same comparison of monthly-mean zonal winds is shown in Figure 6. In these model monthly-means, we are primarily seeing the projections of the slowly varying equatorial Rossby and Kelvin wave modes on the tropopause temperature and wind structure. The stronger asymmetry in rain and

213 heating across the equator in the 2006 El Niño period leads to stronger responses in the equatorial
214 Rossby waves and characteristic off-equator temperature minima. Conversely in 2007, the western
215 Pacific temperature minimum and zonal winds are stronger on the equator and likely also related to
216 the stronger MJO activity there and stronger Kelvin wave responses. Similar tropopause responses
217 to latent heating variations were reported in the idealized model study of Norton [2006].

218 While the model shows differences from MERRA in these monthly-mean comparisons, the de-
219 gree to which our highly idealized model does capture the observed zonally-asymmetric wind and
220 temperature pattern differences in these two years is due to the realism of the monthly-mean heat-
221 ing distribution that is forcing the model. The comparison highlights the importance of waves
222 forced by tropical latent heating in controlling the upper-level circulation and temperature struc-
223 ture.

224 *b. Gravity Wave Spectrum*

225 With heating varying on a 30-min timescale, the response in the model includes a broad spec-
226 trum of gravity waves. We seek to identify relationships between the gravity waves generated
227 by the different latent heating variations, as well as their effects on the circulation in the lower
228 stratosphere. The differences in the zonal mean winds in December 2006 and December 2007
229 are strong at QBO altitudes, beginning at 18 km and above (Fig. 2c). The QBO wind variations
230 will dramatically alter the spectrum of waves through wave-mean flow interaction. We therefore
231 begin here by examining the vertical flux of horizontal momentum, which describes the gravity
232 wave contribution to the EP-flux [Andrews et al. 1987], and we examine this at the tropopause
233 (~ 17 km), which is above the direct latent heating forcing but below the QBO wind influences on
234 the spectrum. This permits an examination of the influences on the tropospheric latent heating and

235 circulation on the vertically propagating wave spectrum in isolation from stratospheric wave-mean
 236 flow interactions.

We perform 3-dimensional spectral analysis as a function of longitude, latitude, and time on overlapping 50° longitudinal sectors spanning latitudes $\pm 25^\circ$ over 3-day periods. Wind anomalies in the zonal and meridional directions (u', v') are computed as deviations from the $50^\circ \times 50^\circ$ sector trends, and cosine taper functions in latitude and longitude are applied. The longitude sectors overlap by 5° on each side where the taper=1/2 such that the total flux in all sector spectra equals the global total, and each spectrum after tapering represents a $\sim 40^\circ \times 40^\circ$ region. The effective horizontal wavelength range resolved in the spectrum is 227-4447 km. Complex 3-dimensional transforms ($\hat{U}, \hat{V}$) are then multiplied by the complex conjugate of the vertical wind transform $\hat{W}^*$ computed on the same grid, and the real part multiplied by atmospheric density gives the spectral density of vertical flux of horizontal momentum

$$F_M = \rho [\text{Re}(\hat{U}\hat{W}^*), \text{Re}(\hat{V}\hat{W}^*)] / \Delta k_x / \Delta k_y / \Delta \omega,$$

237 where (k_x, k_y) is the horizontal wavenumber vector and ω the frequency. The results are rebinned in
 238 terms of azimuthal direction of propagation ($^\circ$ from east) and phase speed (c) and renormalized to
 239 spectral density in these coordinates. Results for Dec 2006 and 2007 are shown in Figure 7, where
 240 we have further reduced the maximum wavelength included in these spectra to wavenumbers > 12
 241 (3335 km). The spectra have been averaged over 15-17 km (above the direct latent heating forcing
 242 and below the QBO) to best represent the gravity waves entering the stratosphere prior to their
 243 interaction with the QBO winds.

244 The spectra show some clear differences, but mainly similarities. 2006 shows a weak preference
 245 for westward propagation compared to eastward, while 2007 shows a strong peak in the east-
 246 northeast direction and relatively weak westward flux. Meridional fluxes are more similar in the

247 two years. At higher phase speeds, $c > 20 \text{ m s}^{-1}$ the spectra are very similar, displaying a much
 248 broader westward spectrum and narrower eastward spectrum. To bring out phase speeds relevant
 249 to the QBO, Figure 8 shows the phase speed spectrum of the zonal flux only. The La Niña case
 250 (2007) has larger flux overall, while the El Niño case shows slightly larger fluxes over a narrow
 251 range near $c = -20 \text{ m s}^{-1}$. Note, however, that we do not examine north-south asymmetries in
 252 gravity waves, which have shown sensitivity to ENSO in previous work [Sato et al. 2016].

253 Figures 9 and 10 examine regional variations in the spectrum. Not surprisingly, these are sub-
 254 stantial, and the fluxes vary to some degree with regional variations in the heating. Clearly the
 255 ENSO variations in heating give rise to strong regional variations in the gravity wave spectra al-
 256 though the zonal mean spectra (Figs. 7 and 8) were relatively more similar. Surprisingly, the strong
 257 El Niño heating in the Dec 2006 central Pacific does not result in much stronger gravity wave mo-
 258 mentum fluxes. The reason is likely related to different tropopause winds (Fig. 9c): Tsuchiya et al.
 259 [2016] found stronger tropical gravity wave activity correlated with westward tropopause winds,
 260 while in the central and eastern Pacific those winds are eastward. Note that the spectra over the
 261 Indian Ocean and S. America show secondary peaks in westward propagating waves at the slowest
 262 phase speeds. This is a spectral signature of the obstacle effect for wave generation associated with
 263 the upper troposphere westward winds interacting with deep convection [Alexander et al. 2006].
 264 If these waves were instead generated in the middle troposphere, they would have been filtered by
 265 the upper troposphere westward winds. These regional spectra also make it clear that the strongest
 266 east/west asymmetries occur over these regions plus the African and S. American tropics, where
 267 we see much faster westward phase speeds and much stronger eastward fluxes at $c < 10 \text{ m s}^{-1}$.

268 Figs. 9c and 10c summarize the geographical and interannual variations in tropopause gravity
 269 wave momentum flux more quantitatively. Each symbol represents absolute momentum flux in a
 270 $40^\circ \times 40^\circ$ region, $\pm 20^\circ$ latitude straddling the equator. In both of our simulated years, the Indian

271 Ocean and Maritime Continent regions are the locus of strongest tropopause gravity waves, where
 272 tropopause winds are also westward. The strongest fluxes in our simulations occur in December
 273 2007 in the sector surrounding Sumatra, a locus of MJO activity throughout an extended portion
 274 of the month (Fig. 2), suggesting that MJO precipitation variability might have greater effect on
 275 gravity wave momentum fluxes than ENSO variability. Note that while a strong peak in total omni-
 276 directional flux occurs in the central Pacific in the 2006 El Niño year, the peak in total zonal flux is
 277 muted due to strong filtering of eastward propagating waves in this region of eastward tropopause
 278 winds. In the 2007 La Niña year we see a peak in the Sumatra sector in total omni-directional
 279 flux as well as a strong peak in zonal flux, which is associated with the peak in eastward waves
 280 (Fig. 9b) propagating through westward tropopause winds.

281 *c. Validation of modeled momentum fluxes*

282 Recent research has highlighted the high degree of intermittency in the occurrence of gravity
 283 waves with different amplitudes [Hertzog et al. 2008; 2012; Jewtoukoff et al. 2013; Plougonven
 284 et al. 2013; Alexander and Grimsdell 2013; Alexander 2015; Wright et al. 2013]. Observations
 285 display log-normal distributions of gravity wave momentum fluxes with infrequent but large events
 286 that contain much of the total flux. This high degree of intermittency can be quite important to
 287 gravity wave effects on circulation [de la Cámara et al. 2014; Bushell et al. 2015]. Specifically,
 288 the larger amplitude waves in the tails of the distributions can break at lower altitudes and result in
 289 larger forcing in the stratosphere. Here we examine mean gravity wave momentum fluxes as well
 290 as momentum flux frequency distributions in the model, and we compare both to observations.

291 Zonal and meridional wind anomalies are computed as in section 4.2 and the products of hor-
 292 izontal and vertical anomalies and density $\frac{1}{2}\rho((u'w')^2 + (v'w')^2)^{1/2}$ give an estimate of the local
 293 momentum flux magnitudes. This method gives accurate maximum values if waves are intermit-

294 tent such that packets appear in relative isolation, an assumption relevant to the lower stratosphere.
295 Ideally, the wind covariances would be averaged over a period or wavelength, however we use this
296 approximate method following previous work [Plougonven et al. 2013].

297 Figure 11a shows occurrence frequencies of these gravity wave momentum fluxes at 20 km, con-
298 trasting Dec 2006 (red) and 2007 (blue), and observations from limb-sounding satellites [Alexan-
299 der 2015] (“HIRDLS/COSMIC”, black). The number of measurements in the satellite retrievals
300 in a single month is too small to fill a distribution, so we use 13 month totals Dec 2006-Dec 2007
301 for comparison to the model results in Fig. 11a. Means of the distributions are 5.2 mPa for the Dec
302 2006 model, and 5.8 mPa for the Dec 2007 model. Averaging the available observations we obtain
303 a mean of 3.2 mPa in Dec 2006 (522 measurements) and 3.9 mPa in Dec 2007 (323 measure-
304 ments). Note that to determine momentum flux from the observations required a wavelet analysis
305 of the vertical structure, so it is not truly a measurement at a single level, but it combines wave
306 amplitude information over a range of altitudes that varies with the wave vertical wavelength. Dec
307 2007 fluxes are larger than Dec 2006 in both the observations and the model.

308 Pre-Concordiasi long-duration balloon measurements covered all longitudes at an altitude near
309 20 km [Jewtoukoff et al. 2013]. The Pre-Concordiasi campaign occurred during a 3 month period
310 February-May, 2010, and the campaign average momentum fluxes from the two tropical balloons
311 were reported at 3.9 and 5.4 mPa. The Pre-Concordiasi values include the spectrum of gravity
312 waves from the buoyancy period to 1-day period, while our model cannot represent waves with
313 periods shorter than 1 hour due to the 30-min resolution of the forcing but includes periods up to 3
314 days. Thus neither the satellite nor the balloon observational comparisons can be considered exact,
315 but these comparisons do suggest the modeled fluxes are reasonably similar to the observations.

316 All of these momentum flux distributions approximately represent log-normal, non-Gaussian
317 distributions. The standard deviation is therefore large compared to the mean value, but this does

318 not indicate a lack of significance. Jewtoukoff et al. [2013] considered the difference between the
 319 two Pre-Concordiasi balloons as significant and attributed the difference to geographic sampling:
 320 The second balloon with larger mean flux spent more time above the Indian Ocean/West Pacific
 321 Ocean sectors where the occurrence of multicellular convection was concentrated and where we
 322 also see largest momentum fluxes. Similarly, it is likely that the differences between the model
 323 in Dec 2006 and Dec 2007 are also significant given the distinct peaks in the fluxes seen in the
 324 spectrum (Fig. 7). However, the uncertainties in the heating derived from CMORPH precipitation
 325 would also need to be considered in order to claim a statistically significant difference between
 326 the two cases. Both the 3D satellite observations and the model suggest somewhat larger fluxes in
 327 Dec 2007 than in Dec 2006.

328 At 20 km, QBO wind shears will have filtered some of the gravity waves. To isolate differences
 329 associated with the tropical tropospheric conditions in the two years, we also show modeled mo-
 330 mentum flux distributions at the tropopause (~ 17 km) in Fig. 11b, an altitude just below the QBO
 331 shear zones. Note the expanded abscissa range to show the higher values that occur at this altitude.
 332 Here interannual differences appear more prominently in the extended tail of the distribution, and
 333 statistics for these distributions are shown in Table 1. The percentile statistics indicate that in Dec
 334 2007 for example, fluxes larger than 20 mPa occur only 10% of the time (90th percentile) but
 335 correspond to 54% of the total flux. This indicates that while convective waves are not quite as
 336 intermittent as orographic waves observed over Southern Hemisphere topography [Hertzog et al.
 337 2012; de la Cámara and Lott 2015], the convective waves display a substantially larger degree of
 338 intermittency than is commonly assumed in non-orographic gravity wave parameterizations (see
 339 Bushell et al. [2015]). By comparing our mean tropopause fluxes to the 20 km values given earlier,
 340 a large fraction ($\sim 45\%$) of the flux has already dissipated in QBO shear zones below 20 km.

d. QBO Wave Driving

We can also examine the tropical wave EP-fluxes and flux divergences in the model. EP-flux divergence is a measure of the wave drag forces acting on the QBO. We also investigate the types of waves responsible for these forces in the model.

Figure 12 shows zonal wavenumber-frequency spectra of the absolute value of the vertical component of the EP-flux at 20 km and profiles of the divergence of this flux in Dec 2006 (a,b) and 2007 (c,d). In b and d, two profiles of each color represent separate integrations over eastward-only (positive) or westward-only (negative) wavenumbers. Black profiles represent integrals over the full halves of the spectrum. Red profiles are integrals over only the planetary-scale waves, which we define as frequencies less than 1 cyc d^{-1} and wavenumbers less than 12, illustrated with a small box near the origin in the two figures. The differences between red and black profiles then show the contributions from eastward and westward propagating gravity waves to the force on the circulation.

In 2006, gravity waves account for almost all the westward forcing, whereas in 2007 the eastward forcing is more equally proportioned between gravity waves and Kelvin waves. Note also that in both cases contributions from gravity waves are substantial even below 20km. Close examination of the two panels in Fig. 11 reveal that changes in the momentum flux distributions between 17 km and 20 km are mostly due to the loss of infrequently occurring, large amplitude waves, and similar changes with altitude have also been seen in other models [Hertzog et al. 2012]. We note that these large amplitude waves are missing in parameterizations of convective gravity wave drag (see Bushell et al. [2015]), which may explain why models tend to poorly represent the QBO in the lower stratosphere. Conversely, our model with realistic distributions in gravity wave sources

(i.e. latent heating) generates a much more realistic distribution of gravity wave amplitudes, and hence significant gravity wave forces in the lowermost stratosphere.

5. Discussion

Recent work has shown clearly the very intermittent nature of gravity waves. The intermittency in our simulations (Fig. 11a) compares well to observations at an altitude near 20 km. Bushell et al. [2015] show tropical momentum flux distributions for different gravity wave parameterizations (their Fig. 6). Their invariant non-orographic parameterization dropped 4 decades in occurrence at a flux of 6 mPa. Essentially, all of the waves in the parameterization are weak in amplitude and not intermittent. They also showed the distribution of gravity wave momentum fluxes using a variable convective source parameterization. In this case, occurrences drop 4 decades at a flux of ~ 20 -25 mPa, which is much more realistically intermittent than the invariant parameterization, but the intermittency falls far short of that observed or that produced in our model. In particular, long-duration balloon observations [Jewtoukoff et al. 2013] (their Fig. 15) show that the momentum fluxes drop 4 decades in occurrence at flux values ~ 100 mPa, and this occurs at ~ 80 mPa in our simulations.

Typical invariant non-orographic gravity wave parameterizations have only very weak forces in the stratosphere. They are designed instead to give realistic circulation effects in the mesosphere. Orographic gravity waves are parameterized with much larger amplitudes than non-orographic waves, and as a result they break and change the circulation in the upper troposphere and lower stratosphere. However, large amplitude waves from convection do occur, and the momentum flux convergences in the stratosphere can lead to substantial forces. For example, Stephan et al. [2016] showed that realistic waves from summertime convection over the U.S. produce forces in the lower stratosphere that rival orographic wave forcing. Most parameterizations in models

386 give relatively very small wave forces at stratospheric levels. Stochastic non-orographic parame-
387 terization methods that account for realistically intermittent amplitudes have been developed [de
388 la Cámara and Lott 2015], and implementation in a global model showed improvements in the
389 timing of the springtime transition from westerly to easterly winds in the Southern Hemisphere
390 stratosphere [de la Cámara et al. 2016]. So including realistic intermittency in parameterized
391 non-orographic gravity wave amplitudes, while simultaneously reducing gravity wave drag due
392 to orographic waves, may be a way forward. Indeed, climate models struggle to simultaneously
393 simulate realistic Northern and Southern Hemispheric stratospheric winds, which could be due to
394 an over-reliance on orographic gravity wave drag.

395 Similarly, we find support for much larger intermittency in tropical convective gravity waves
396 than is typically parameterized, and we hypothesize this is the reason that models struggle to
397 represent realistic QBO winds and wind shears in the lower stratosphere at levels below 40 hPa
398 (e.g. Krismer and Giorgetta [2014]; Richter et al. [2014]; Coy et al. [2016]). Typical gravity
399 wave parameterizations drive only the upper levels of the QBO while planetary scale waves are
400 responsible for most or all of the forcing at the lower levels. An early example with an invariant
401 parameterization was shown in Giorgetta et al. [2002]. More recently Richter et al. [2014] showed
402 modern results with a variable convective source parameterization that gave a very realistic QBO
403 at pressure levels above 40 hPa, but in the lower stratosphere the westerly phases are too strong and
404 easterly phases too weak. Yoo and Son [2016] have shown that easterly QBO winds in the lower
405 stratosphere are associated with stronger tropical intraseasonal precipitation in the observational
406 record. Hence such errors in modeled QBO winds may hinder a model’s ability to represent the
407 observed stratosphere-troposphere connections. We also note that many previous studies have
408 suggested that the easterly QBO wind phases are forced primarily by gravity wave drag (e.g.
409 Dunkerton [1997]; Kawatani et al. [2010a]).

410 More realistic intermittency such as shown in our Fig. 11 does in fact lead to significant forces
411 in the lower stratosphere below 20 km (Fig. 12). That these forces are due to dissipation of the
412 largest amplitude waves is also evident from comparison of the distributions at 17 km and 20 km
413 shown in Fig. 11. Nearly half of the gravity wave momentum flux is dissipated between these
414 levels in our model.

415 Our results may be relevant for realizing the long-range forecast skill that is expected from re-
416 alistic representation of the tropical stratosphere in forecast models. Although the Scaife et al.
417 [2014a] study found the QBO among the four leading sources of skill in their winter seasonal
418 forecasts of the NAO, their forecast model's QBO teleconnection pattern was weaker than in the
419 observations. As mentioned above, model representations of the QBO tend to be least realistic at
420 low levels below 40 hPa, and discrepancies in the width of the QBO are also common [OSullivan
421 and Young 1992; Hansen et al. 2013]: Either or both of these could be reasons for weaker tele-
422 connections in models. Maximum correlations between extratropical winter conditions and QBO
423 winds have been observed with 50 hPa QBO wind in observations. If the lower levels of the QBO
424 are unduly important to describing extratropical teleconnection strength, it points to a clear weak-
425 ness in models. Further, the results of Yoo and Son [2016] suggest that long-range forecasting
426 skill in tropical intraseasonal precipitation may be tied to realistic representation of the QBO at
427 lower stratosphere levels in models.

428 While studies have shown the QBO to be highly predictable on time scales longer than a year
429 [Scaife et al. 2014b] the unprecedented disruption of the QBO in 2016 and the failure of forecast
430 models to predict it [Newman et al. 2016; Osprey et al. 2016] place new emphasis on more realistic
431 representation of the wave forcing of the QBO. There is also observational evidence that the QBO
432 winds at low levels near 70 hPa may be experiencing a long-term weakening trend [Kawatani and

Hamilton 2013]. Hence more realistic simulation of the QBO may also be beneficial to near-term climate prediction as well as seasonal forecast model skill.

In addition to forcing the stratosphere and mesosphere, gravity waves from convection can also directly force the circulation in the upper atmosphere and ionosphere [Vadas and Liu 2013; Vadas et al. 2014]. The gravity waves that can propagate to these high altitudes have fast phase speeds, faster than $\sim 50 \text{ m s}^{-1}$. While the peaks in our integrated phase spectra (Figs. 7-8) occurred at phase speeds $\sim 7\text{-}20 \text{ m s}^{-1}$, the spectra in Fig. 12 show that much faster waves also appear at higher frequencies. In These maps also make it clear that the strongest east/west asymmetries occur over these regions plus the African and S. American tropics, where we see much faster westward phase speeds and much stronger eastward fluxes at $c < 10 \text{ m s}^{-1}$. particular, a lobe with phase speeds of 70 m s^{-1} among the westward propagating highest frequencies is prominent. According to the linear dispersion relation (neglecting wind effects) vertical wavelength $\lambda_z \sim 2\pi c/N$, these fast waves would have $\lambda_z \sim 44 \text{ km}$ in the troposphere, which is close to four times the most common cloud and heating depth in our simulations of 11 km (Fig. 3). While a vertical wavelength of twice the depth of the heating, or 22 km , is predicted for large-scale heat sources, Holton et al. [2002] showed that smaller-scale heat sources will project more strongly on vertical wavelengths four times the depth of the heating. Our model simulations support the Holton et al. [2002] result, and show that such fast waves clearly appear in our simulations. In fact, they can dominate the convectively-generated gravity wave spectrum at wave periods shorter than a few hours.

6. Summary and Conclusions

We use satellite-based global precipitation and cloud data at high spatial and temporal resolution to estimate three-dimensional time-varying latent heating and the resulting global wave spectrum generated by convection. The modeled zonally-averaged gravity wave momentum fluxes in the

456 lower stratosphere are similar to those derived from 3D satellite data, and similar to those observed
457 by Pre-Concordiasi long-duration balloons. Modeled distributions of gravity wave momentum
458 fluxes also display similar intermittency to the Pre-Concordiasi balloon measurements. These
459 comparisons show that the modeled zonally-averaged fluxes fall within the range of variability
460 seen in observations.

461 Interannual variations in gravity waves were examined in the context of interannual precipitation
462 variations characterized by the ENSO 3.4 index. Spectra and intermittency of momentum flux
463 were also evaluated. Profiles of momentum flux convergence were used to examine gravity wave
464 forces acting on the QBO shear zones, and these forces were compared to planetary-scale tropical
465 wave forces. The results show that in the zonal mean sense, the changes with ENSO are only
466 modest, although regional variations in the gravity waves are large. For example, despite more
467 rain and latent heating in the El Niño case, the zonal gravity wave momentum fluxes are smaller
468 than in the La Niña case because of the shift in the precipitation to the central Pacific where upper
469 tropospheric zonal winds are less favorable for vertical wave propagation. The more active MJO
470 convection in the Indian Ocean/Maritime Continent region in the La Niña case appears to be a
471 more important source in terms of gravity wave momentum fluxes.

472 The modeled intermittency in gravity wave amplitudes is similar to that observed in existing
473 drifting isopycnal balloon measurements [Jewtoukoff et al. 2013], but current parameterization
474 methods significantly underestimate this degree of intermittency in gravity waves above tropical
475 convection, even with more realistic convective source parameterizations. Stochastic parameteri-
476 zation methods such as described in de la Cámara and Lott [2015] could be applied to the tropics
477 utilizing these intermittency statistics, and we show evidence to suggest that such intermittency
478 could improve the simulation of the QBO at lower levels where models show clear weaknesses,
479 below ~ 22 km (40 hPa). We further hypothesize that improving the simulation of the QBO at these

480 lower altitudes might improve simulation of tropical-extratropical teleconnections and associated
481 skill in long-range weather and seasonal climate forecasts.

482 In the future, we may have better observations to validate the inter-annual and regional variations
483 in gravity wave momentum flux predicted in our model. Future measurements planned during the
484 STRATEOLE-2 field campaign (www.strateole2.org) will provide a wealth of observations for
485 model validation. Beginning in 2014, new precipitation measurements in the Global Precipitation
486 Measurement (GPM) era have led to a new 30-min, $0.1^\circ \times 0.1^\circ$ resolution IMERG rain rate product
487 [Huffmann et al., 2015]. These data are reportedly better constrained at higher frequencies, and
488 may provide more accurate forcing for future model studies that can be more thoroughly validated
489 with observations from STRATEOLE-2.

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492 APPENDIX

493 **Latent Heating**

494 The latent heating algorithm we use to compute space-time gridded heating rates suitable for
495 wave studies was described in Ryu et al. [2011]. There, they showed the zonal-mean heating
496 profiles as functions of latitude in comparison to version 1 of the TRMM CSH latent heating
497 product. Changes in version 2 of CSH resulted in stronger rates and a shift downward in the
498 altitude of the peak heating. (See Tao et al. [2010]: their Fig. 10.) Considering these changes,
499 our heating algorithm compared reasonably well in the mean to CSH. No high-frequency latent
500 heating products exist for us to compare the higher-frequency variability. We instead validate our
501 modeled gravity waves with observations in section 4.

502 A further examination of the heating input to the model is shown in Figure A1. These are average
503 heating profiles over land and ocean regions within $\pm 30^\circ$ latitude for Dec 2006 and Dec 2007. The
504 heating profile shapes compare well to the TRMM SLH 15-yr means over land shown in Liu et
505 al. [2015], although these ocean profiles display weaker secondary shallow heating than the 15-yr
506 SLH means. The El Niño year (Dec 2006) shows less difference between heating over land and
507 ocean than the La Niña year (Dec 2007), which is not surprising given the shifts in precipitation
508 evident from Fig. 4. The active MJO during both of these months may be responsible for the
509 higher peak heating magnitudes in these cases compared to multi-year means [Tao et al. 2010; Liu
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688 10.1002/2016GL067762.

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TABLE 1. Tropopause Momentum Flux Distribution Statistics

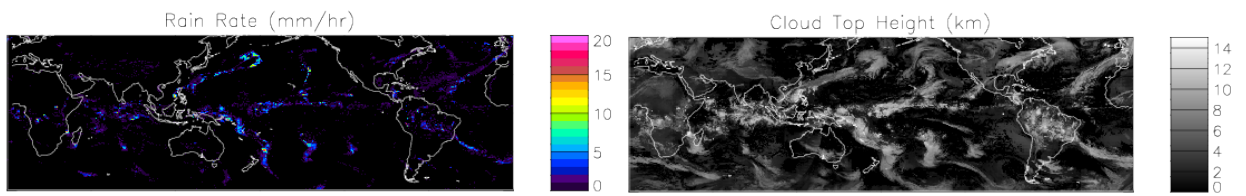
	Mean	90 th	99 th
	(mPa)	Percentile	Percentile
Dec 2006	9.5	19mPa/49%	46mPa/14%
Dec 2007	11	20mPa/54%	51mPa/18%

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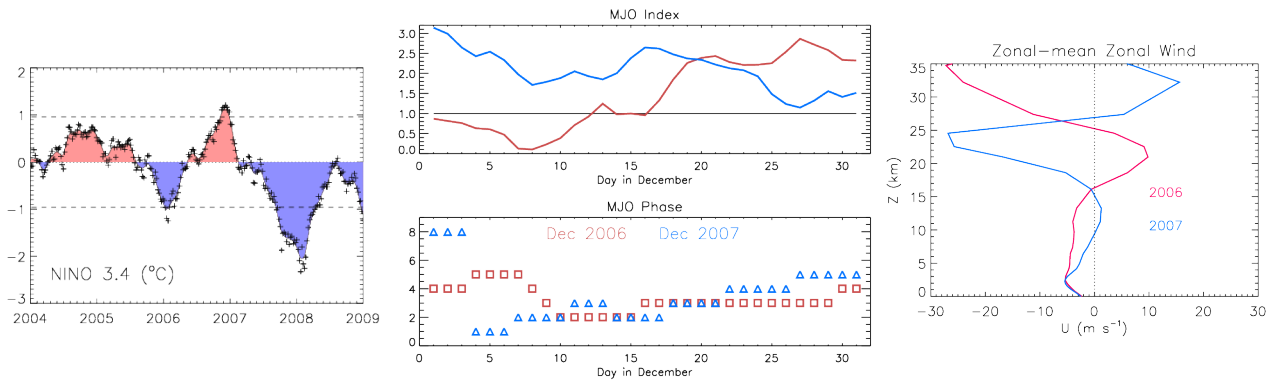
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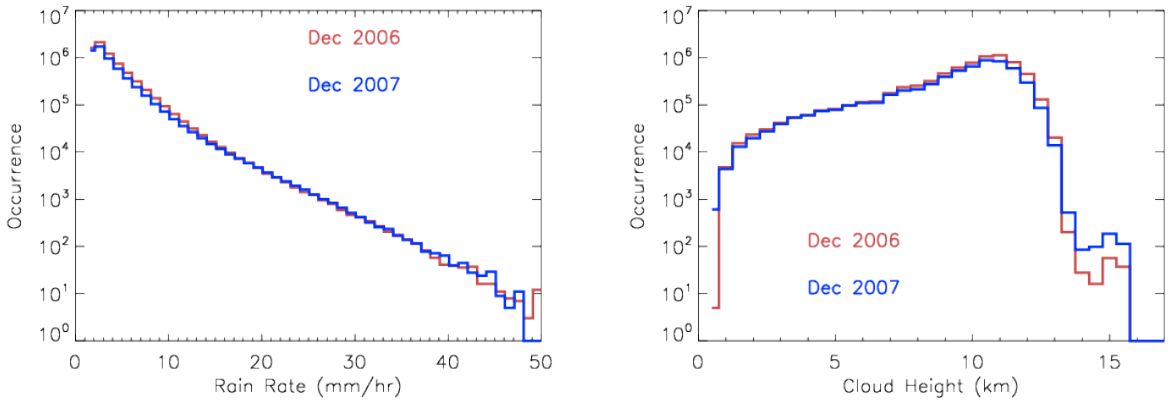
Fig. A1. Profiles of latent heating averaged over land (solid) and ocean (dashed). Left: December 2006. Right: December 2007. 49



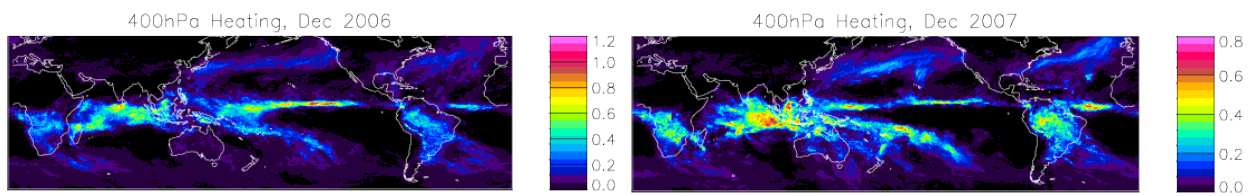
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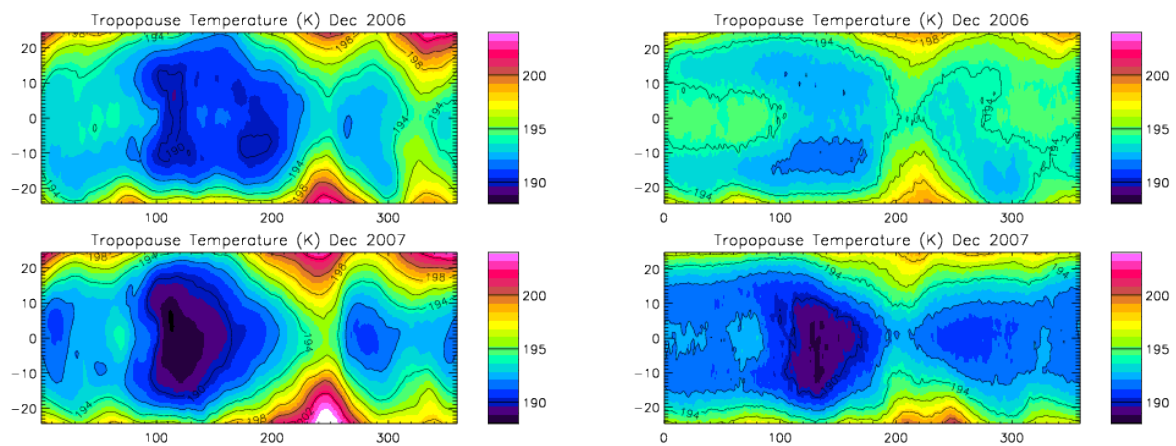
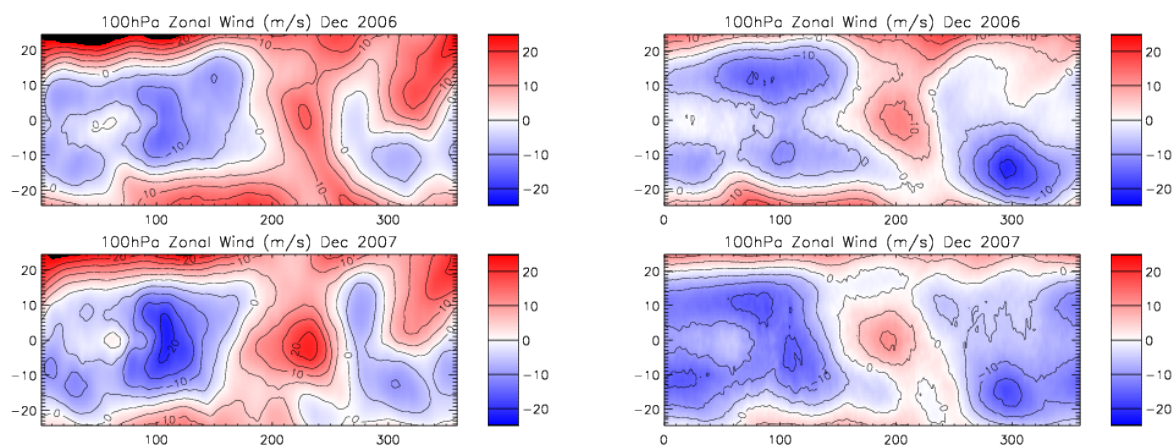
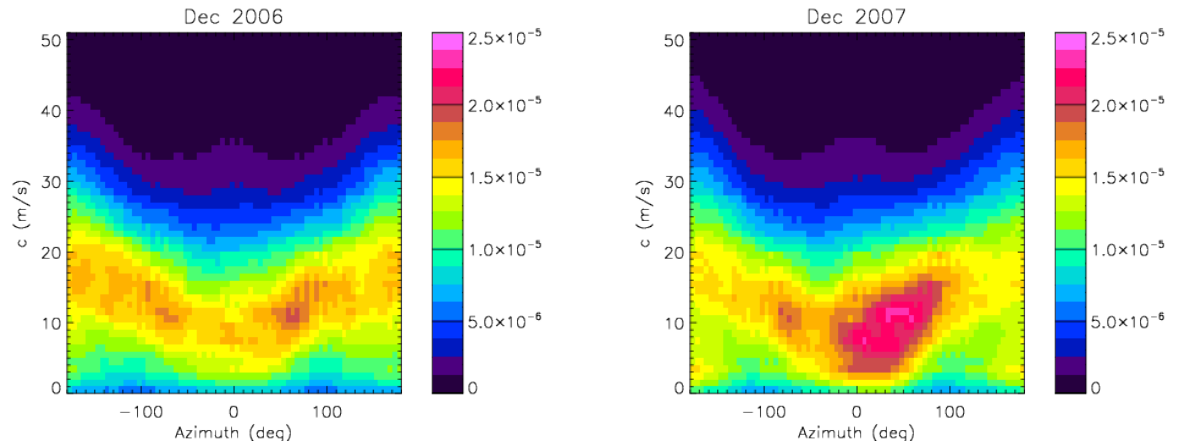


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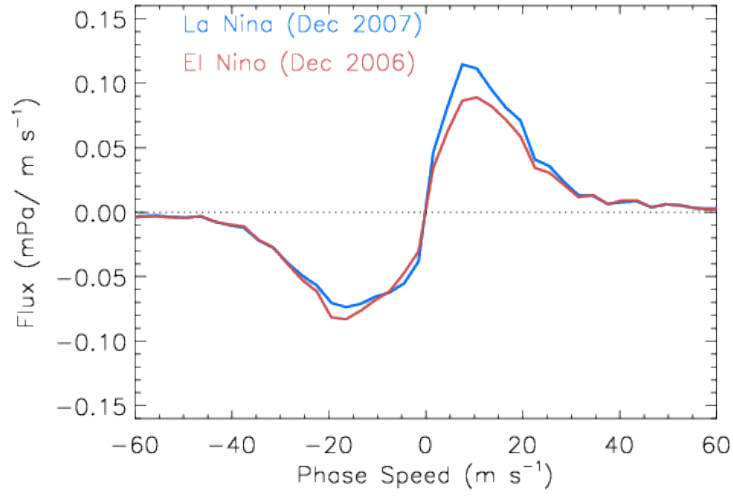


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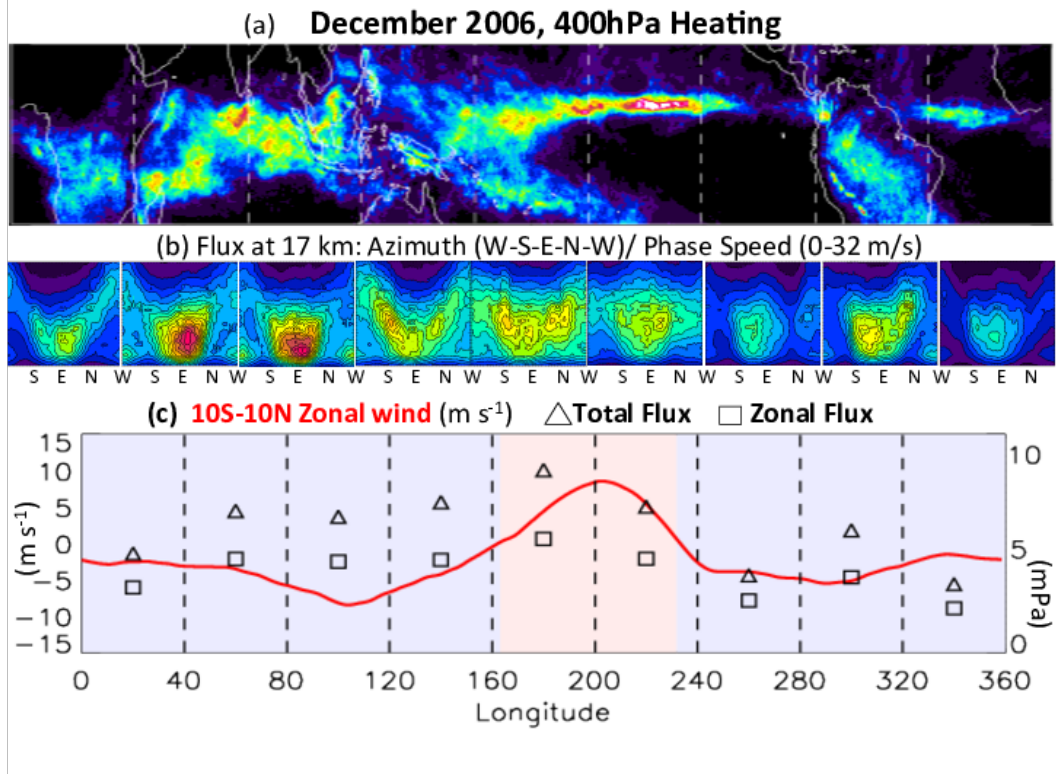


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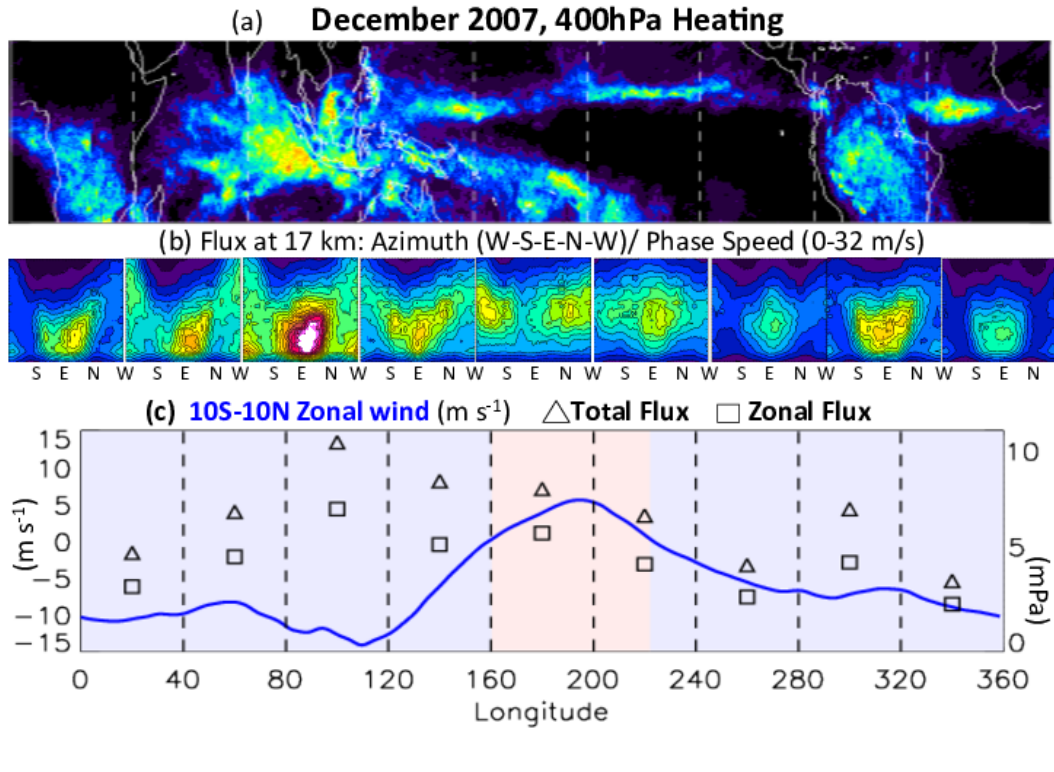
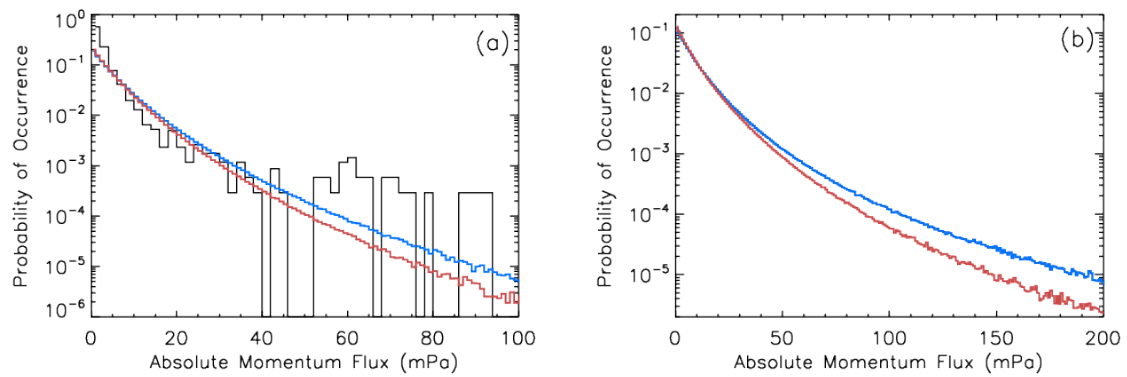


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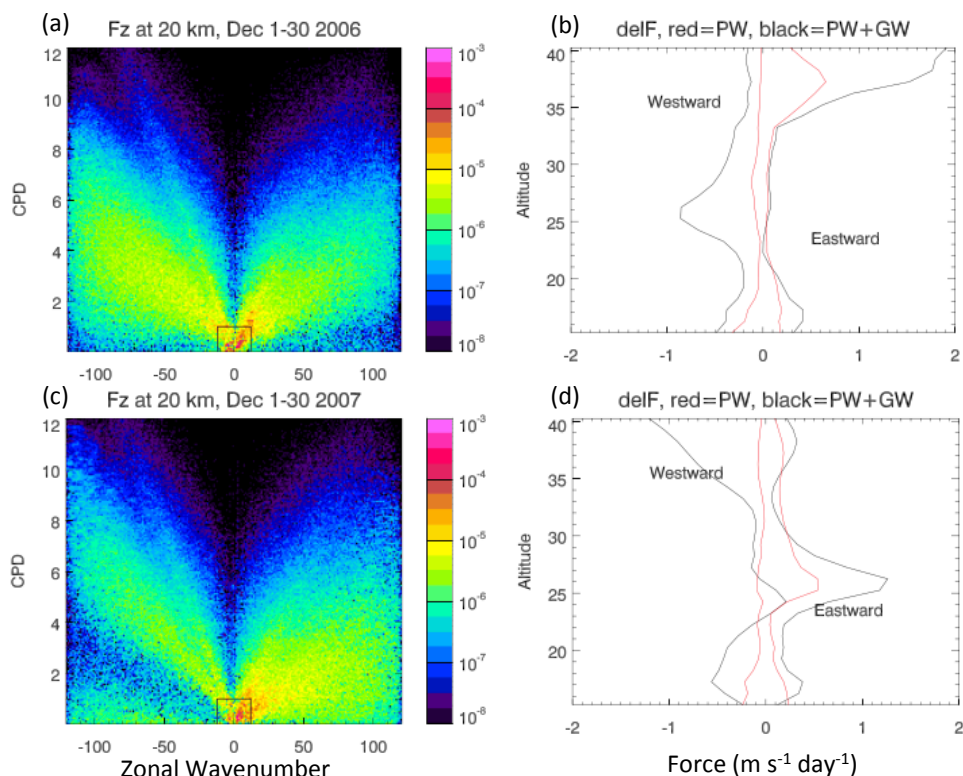


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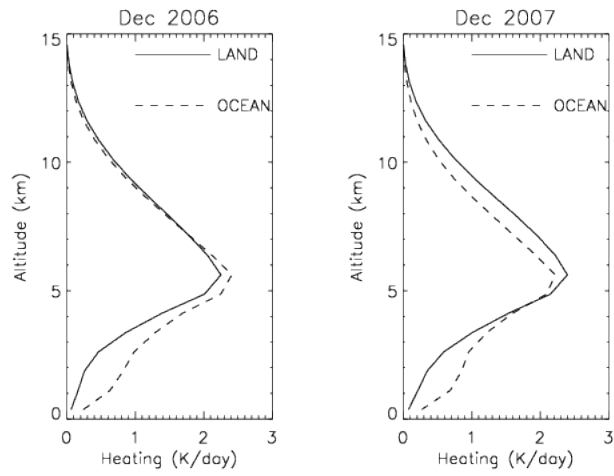


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