• SOLIS (NSO)

Upcoming synoptic vector field data sources:

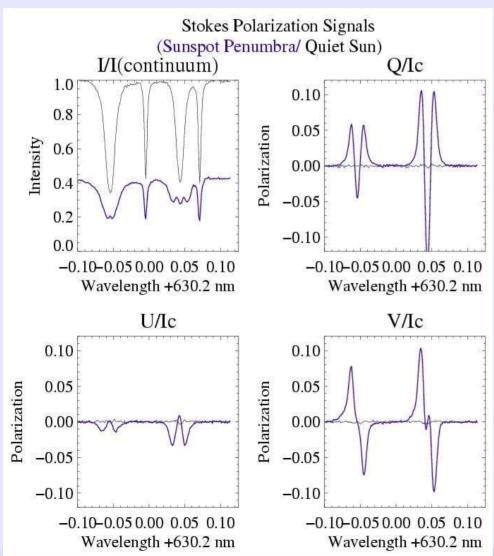
- line-of-sight available now, replaces old KPNO 'grams.
- 1" sampling.
- Vector available soon (any day now ... really).
- Pros:
 - Full disk, benchmark 3/day.
 - Spectrograph-based with full M.E. Inversion: very quantitatively reliable and robust (caveat below).
- Cons:
 - Ground-based, so seeing-influenced results.
 - single-station (for now): low temporal sampling.
- Solar-B (NASA/Japan)
 - Launch scheduled for 2006.
 - Pros:
 - Space-based: no seeing influences.
 - Cons:
 - Small FOV instrument w/ very high resolution spectrograph instrument that uses a small FOV, *plus* a filter-based instrument. Max 2.5' square.
 - Data access/availability TBD
 - Unclear observing sequence/cadence with larger-FOV instrument.
 - Unclear how quantitative the filter-based data will be.

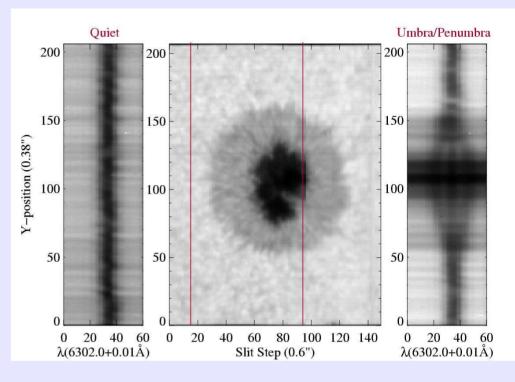
• SDO/HMI (NASA)

- Launch scheduled for 2008
- 1" sampling.
- Pros:
 - Full disk every 90 seconds, space-based.
 - Fully accessible data.
- Cons:
 - Not available until 2008, possibly early 2009.

Measuring the photospheric magnetic field: Stokes spectropolarimetry:

 Zeeman effect: magnetic field induces both energy-level splitting and polarization to emergent light of magnetically sensitive lines.





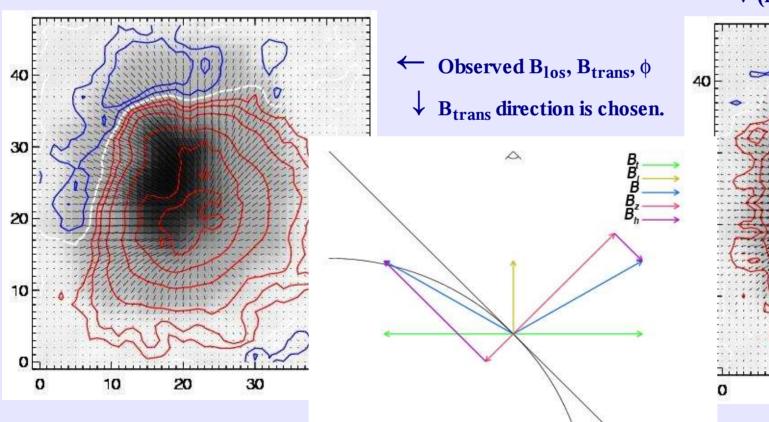
• Shape of polarization spectra and degree of polarization due to: strength, direction of magnetic field, thermodynamics of plasma, spatial and spectral resolution.

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Measuring the photospheric magnetic field cont'd:

- Inversion procedure: I, Q, U, V spectra \rightarrow B_{los}, B_{trans}, ϕ
 - Myriad methods exist, each with strengths and weaknesses
 - Inversion based on Milne-Eddington atmosphere, accounting for Faraday rotation, thermodynamics, magnetic fill-fraction.
- 180° Ambiguity in B_{trans} inherent in Zeeman-polarization observations
 - Myriad methods exist, each with strengths and weaknesses
 - Minimize divergence and current simutaneously
 - Results transformed to heliographic B

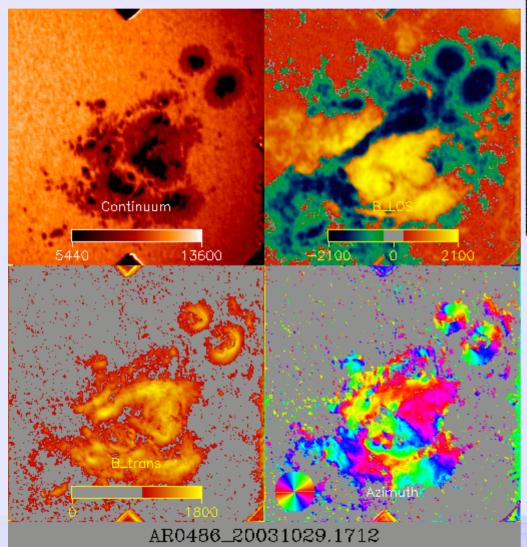
10 20

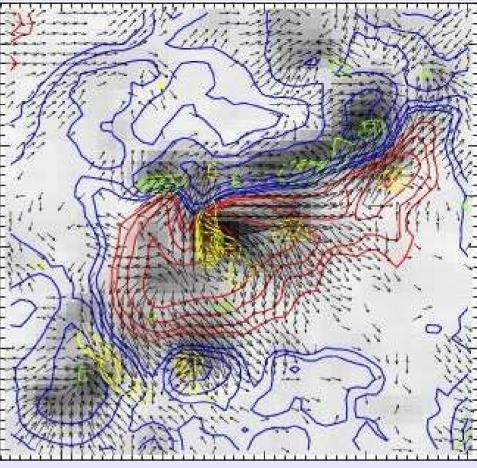


Heliographic B results.

↓ (note shift in neutral line)

Visualization:





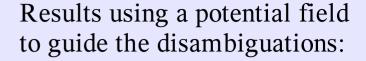
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Ambiguity Resolution: recent workshop approach and example results.

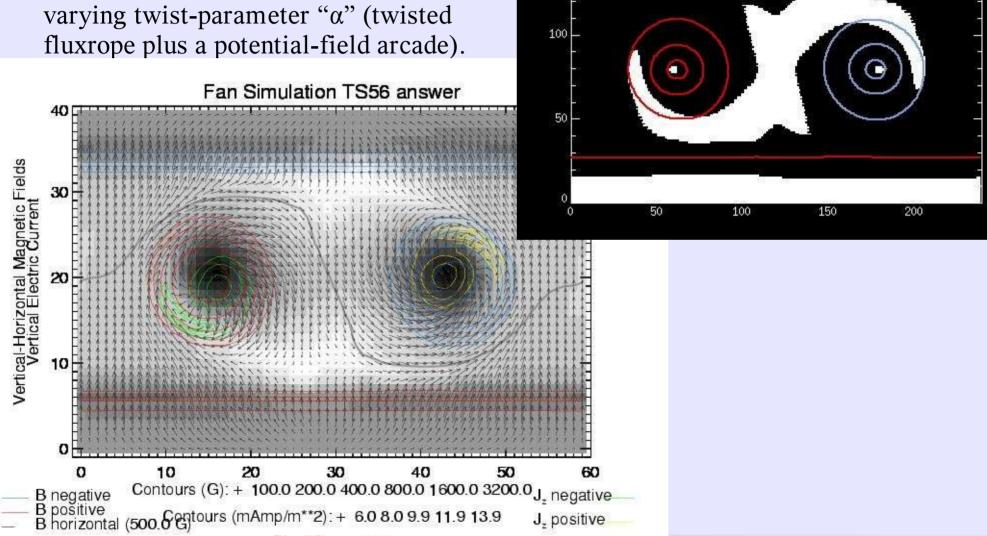
Y. Fan numerical simulation used to construct a "photospheric magnetogram".

Solar-appropriate in many ways.

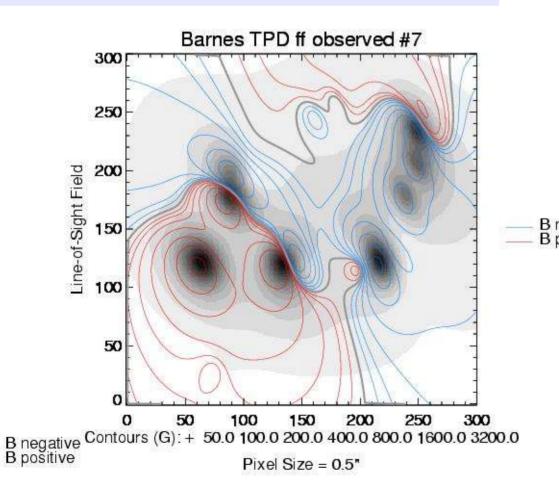
Challenge: a bald-patch, plus widely

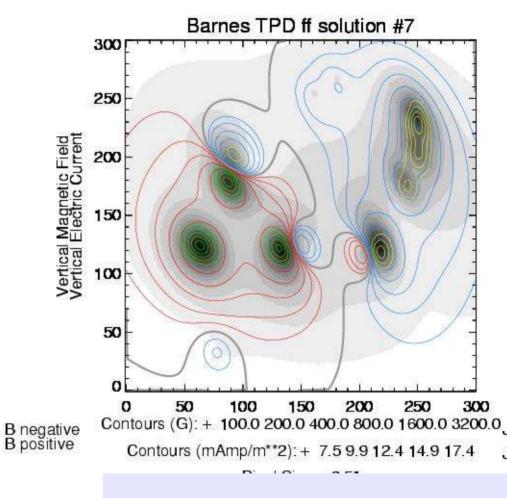


KLeka_Potential_fan_simu_ts56_J.sav



Barnes used approach of Chou & Low to construct a force-free model and placed it at 45 deg west, 20 deg north. Multiple bald-patches, but constant- α . We were afraid this was going to be too easy.





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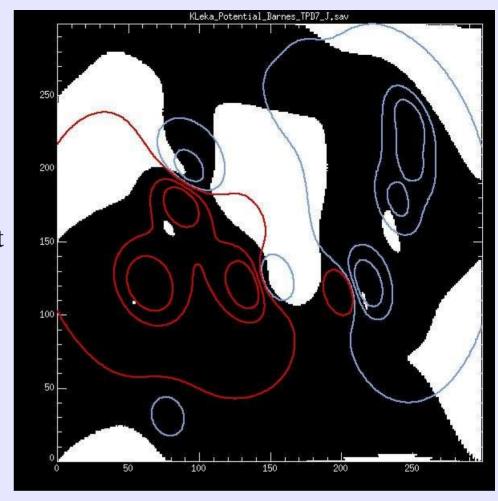
Nope. Results using a potential-field to guide the disambiguation:

Metrics were devised to compare how algorithms performed, based on % area correct, % flux correct, and comparison of vertical currents (where it got it correct *plus* a penalty for extra currents introduced)

Bad news: some algorithms did worse than simply using a potential-field solution.

Good news: a few algorithms that are almost completely autonomous performed very well, scoring > 99% for flux/area metrics, and < 0.1 for the Jz-comparison metric.

Why does anyone care? Need disambiguation algorithm to interpret **B** in a meaningful way; need it automatic for synoptic data such as the HMI firehose.



Discussion Topics:

- Sensitivity: where there is information (active regions), and where there is *not* (under filaments/coronal holes, *etc*)
- Image-coordinates vs. Heliographic coordinates and what that means for vector magnetic field data.
 - "fake" neutral lines and other systematic errors, how it differs w/ vector data vs. MDI-type 'grams.
 - Beware the fine print in research papers.
- How to best present vector **B** data for forecasters
 - Map of a gazillion arrows probably not useful.
 - Parameterizations, quantitative reports for each active region.
 - Tools need to be in place soon. SOLIS data will be available soon and will be a good "prep" for HMI-type data.